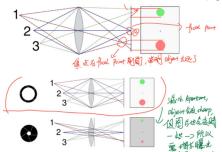
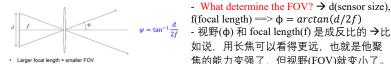


Depth of field(DoF): the distance between nearest and farthest object in a scene that appear acceptably sharp in an image. ==> f-number 越大(e.g., f/22), aperture 越小, 及 DOF 越大, 适合拍自然风光; f-number 越小(e.g., f/1.4), aperature 越大, 及 DOF 越



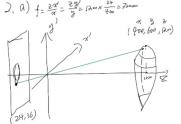
Changing the aperture size affects depth of field Hig (exposure) image Field of view(FOV): how much/wide of scene we can capture?



Dolly zoom: Continuously adjusting the focal length while the camera moves away from (or towards) the subject \rightarrow that's how to get the "Vertigo shot".

Types of Lens aberrations(光差) or flaws related to lens: Vignetting, Radical Distortion--caused by imperfect lenses, Spherical aberration, Chromatic and monochromatic aberration--different refractive indices can cause color fringing, Astigmatism(散光)

Pin hole camera model: (use midterm practice as example)



the world gets imaged or landed in our camera? - How to compute the focus length for a given

小,适合拍近景或人物。

- For a fixed focal length(f) and

contrary, if too close \rightarrow object

- Effect of aperture: smaller

exposure (with tripod) = Sharp

aperture(e.g., f/22) + longer

- What determine the FOV? \rightarrow d(sensor size),

如说,用长焦可以看得更远,也就是他聚

焦的能力变强了,但视野(FOV)就变小了。

will get blurry

haven't converged yet, and image

dist of image plane(D'), if object is

too far away(D) \rightarrow focal points will

converge before image plane; on the

scene point(400, 600, 1200) and image point (24, 36)? \rightarrow f = zx'/ x or zy'/ y

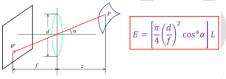
- How did we get $f = \frac{zx'}{x}? \rightarrow$ we are deriving this from x'/x = D'/D, where we assume D = z, and D'=f \rightarrow That's when the distance from lens to camera equals to focal length.

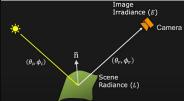
Color filter arrays: Why demosaicing 去马赛 (filling missing pixel value) bias on green? \rightarrow human is more sensitive to green!

Misc. Digital camera artifacts: Sensor noise, in-camera processing(e.g., over-sharpening can produce halos 光晕), compression algorithm (e.g., JPEG artifacts), Blooming (related to CCD charge overflowing), Color artifacts (e.g., purple fringing from micro lenses).

lec3 light and shading:

Image irradiance(E) and scene radiance(L): Image irradiance (E) is proportional to scene radiance(L), area of lens ($\pi d^2/4$), and inverse proportional to (f^2) , and view of field(α)





Reflectance properties:

- Specular reflection: light is reflected about surface normal
- Diffuse reflection: light scatter equally in all direction.
- Other possible effect: transparency, refraction, subsurface scattering, Fluorescence, phosphorescence.
- Why does specular reflection make it more difficult to recognize objects in a scene? \rightarrow Because light is reflected about surface normal, so the viewer can only see the reflected ray along the path of the reflected ray.

Bidirectional reflectance distribution function (BRDF):

- Idea: How bright a surface appears when viewed from one direction when light falls on it from another.
- There are two directions are important to us: 1) the incident direction, where the <u>lights arrived (E)</u>, 2) the emitted direction, in which light is being reflected(L) \rightarrow Definition: the ratio between two.

 $E(\theta_i, \phi_i)$: Irradiance due to source in direction (θ_i, ϕ_i) $L(\theta_r, \phi_r)$: Radiance of surface in direction (θ_r, ϕ_r)

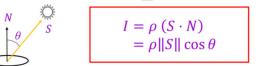
BRDF:
$$f(\theta_i, \phi_i, \theta_r, \phi_r) = \frac{L(\theta_r, \phi_r)}{E(\theta_i, \phi_i)}$$

- isotropic surface: A point in the surface, if the change of rotation of the surface doesn't affect the brightness of points, then that surface is isotropic \rightarrow like diffuse reflection \rightarrow In this case light scatters equally in all direction, so BRDF is constant.

Diffuse reflection: Lambert's law

Unit: 1/sr

Definition: A Lambertian surface has a constant BRDF, which means that a Lambertian surface scatters the entire incident light uniformly in all directions, such that same amount of energy is seem from any direction. Since its BRDF is independent of outgoing directions, Lambertian surface is also called ideal matte or diffusion surface.





I: the reflection intensity, how much light being produced \rightarrow Known

 ρ : albedo, the amount of light being reflected, range bet [0, 1], where 1 means 100% of ray being reflected, and 0 means no reflection. \rightarrow Unknown

S: direction of light source \rightarrow Unknown in general, bec we can't tell if the light is being reflected by object, or object itself is the original light source?

N: surface normal, or unit vector \rightarrow Use to detect whether it's 90deg or not \rightarrow Unknow Q1: Can we recover the surface normal from single image? \rightarrow Only if the surface normal is makes a 0 deg angle with the direction of light source \rightarrow However, in general case, we would need at least three images.

Photometric stereo

Problem statement: Aka shape from shading. Assuming the object is a Lambertian object, and we have more than one photo of an object from multiple light source S_i , and given the image pixel value, I(x, y), can we reconstruct the object shape (N) and albedo $\rho(x, y). \Rightarrow I_i(x, y) = (\rho(x, y)N(x, y)) \cdot (kS_i) = g(x, y) \cdot V_i$

Assumption: 1) A Lambertian object; 2) A local shading model; 3)A set of known light sources;4)a set of pictures of object captured by same camera;5)orthographic projection

lec4: color

Grayscale pixel intensity: These pixel values denote the intensity of the pixels. For a grayscale or b&w image, we have pixel values ranging from 0 to 255. The smaller numbers closer to zero represent the darker shade while the larger numbers closer to 255 represent the lighter or the white color. mage file

Digital Image File Formats

Example of PPM image:

- The header P3 indicates it is a color image,
- with size 3 x 3, and the maximum value is 255

The data is ordered from top to bottom, and left to right. The ordering is RGB. HSV Color Space: think of HSV as a transformation of an RGB colorspace.

When would we use HSV over RGB?

0 0 0 255 0 0 0 0 0 0 255 0 0 0 0 255 255 0 0 0 0 0 0 0 255 0 0 0

P3 3 3 255

• near G: $H = \frac{\pi}{3} \frac{B-R}{V-min(R,G,B)} + \frac{2\pi}{3}$
• near B: $H=rac{\pi}{3}rac{R-G}{V-min(R,G,B)}+rac{4\pi}{3}$
• near R: $H=rac{\pi}{3}rac{G-B}{V-min(R,G,B)}+2\pi$
Value calculation:
V=(R+B+G)/3 or $V=max(R,G,B)$
Saturation calculation:
S = (V - min(R,G,B))/V

HSV is more robust towards external lighting changes. This means that in cases of minor changes in external lighting (such as pale shadows, etc.) Hue values vary relatively lesser than RGB values. Besides, it's fast to convert. \rightarrow But, in general it's not as good as RGB colorspace

Problem of Object Detection: Assuming we have a grayscale image, how do we find bounding box around black object in grayscale background? Algo1-- Black color & Projections:

- Idea: 1) Count number of black pixels in each row and column; 2) analyze these histograms or projections of black pixels onto x- and y-axes
- Advantage: 1) tell us more about the shape of object, and how prevalent a particular value a pixel is. 2) By further analyzing the pattern of histogram, we can identify where the object is.

Algo2--Flood fill (alternative of algo1):

8					
-1	-1	1	1	-1	
-1	-1	1	1	-1	
1	1	-1	1	1	
1	1	<u>-1</u>	1	1	
1	1	1	1	41	

1. Convert to binary image, use -1 or 0 to represent black pixels and 1 to represent white pixels 2. scan row by row until first black pixel is reached and label it object 2 (Object label starts from 2 to differentiate from white pixels) 3. Find all neighbors of the current pixel that are black and assign the object label of the current object in a

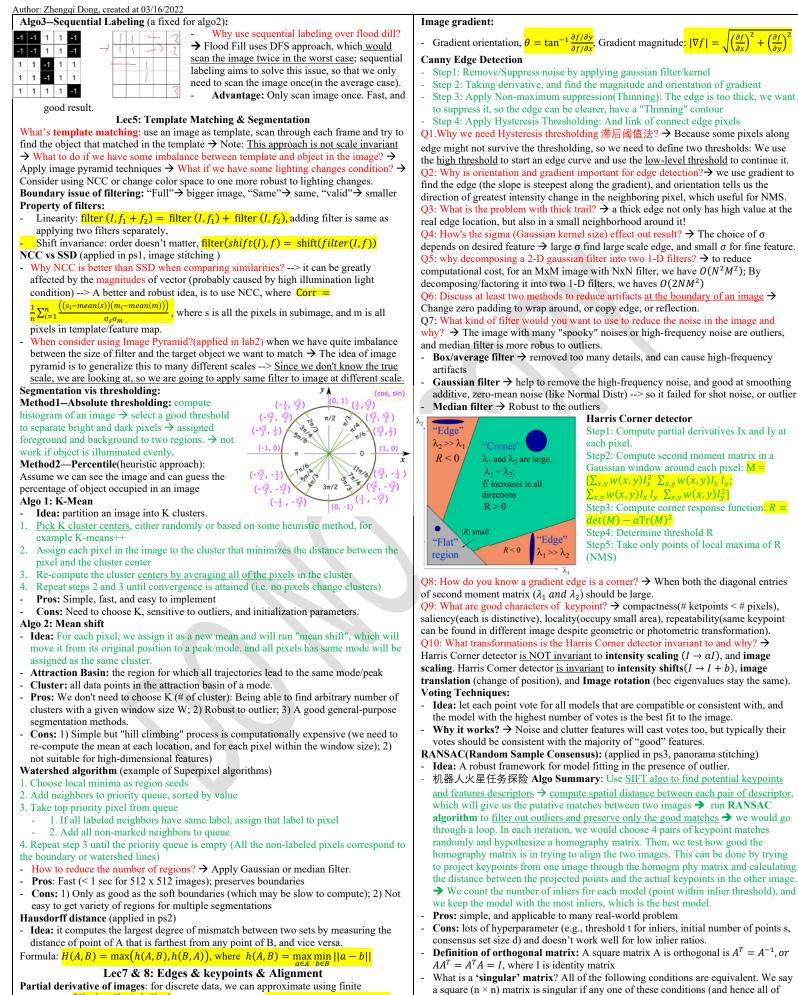
recursive, DFS manner 4. When there are no more black neighbors, continue scanning the image until next black pixel is reached and label it with the next object label and go to step 3.

- Assume: P is a point a surface; P' is the point at image plane.
- Scene radiance(L): 发光度,
- lights emitted from surface P to P'. It measured how much energy leaved surface or being reflected (Watts per

sq. meter per steradian) \rightarrow Note: P can be the original light source, or can be an item that is reflecting light.

- Image irradiance(E): 光照度, lights fall on image plane P'. The image irradiance of a point in the image plane is defined to be the power per unit area of radiant energy falling on the image plane. Radiance is outgoing energy; irradiance is incoming energy.

- We want to know where a point (X, Y, Z) in



Partial derivative of images: for discrete data, we can approximate using finite differences: $\frac{\partial f(x,y)}{\partial x} = \frac{f(x+1,y)-f(x,y)}{\partial x}$ \rightarrow If conv with filter [f(x, y) f(x+1, y)].T @ [-1, 1]

them) is satisfied: 1) det(A) = 0; 2) some columns or columns are not linearly independent; 3) the matrix is not invertible; 4) the matrix is not full rank (rank < n).